Perfect fluid flow from granular jet impact

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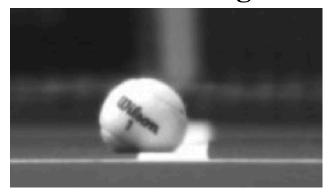
Physics Department & James Franck Institute
University of Chicago

Institute for Computing Science (ICis)
Verification, Validation and Uncertainty Quantification
across Disciplines
Park City, Utah August 2011

Overview

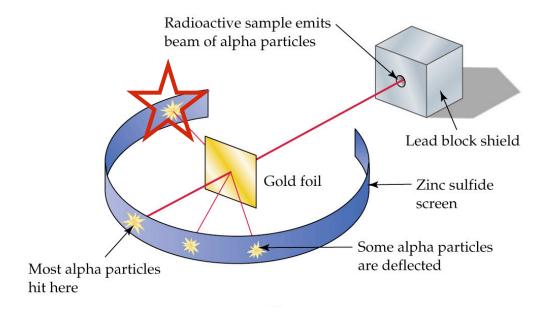
Impact is familiar, important and a powerful experimental tool

mechanics tennis ball hitting court line



Hawkeye innovations

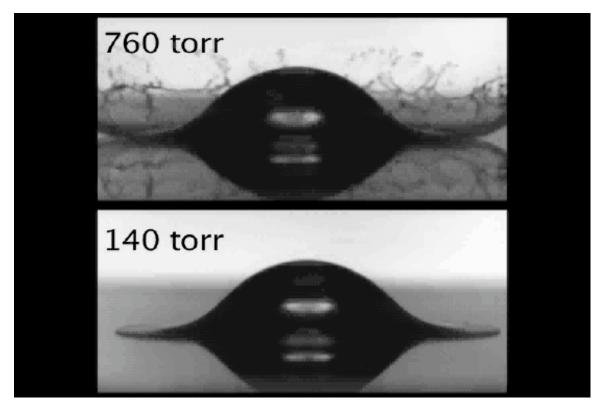
high-energy experimental physics impact → scattering → structure



Rutherford's goldfoil scattering experiment wikipedia

Overview

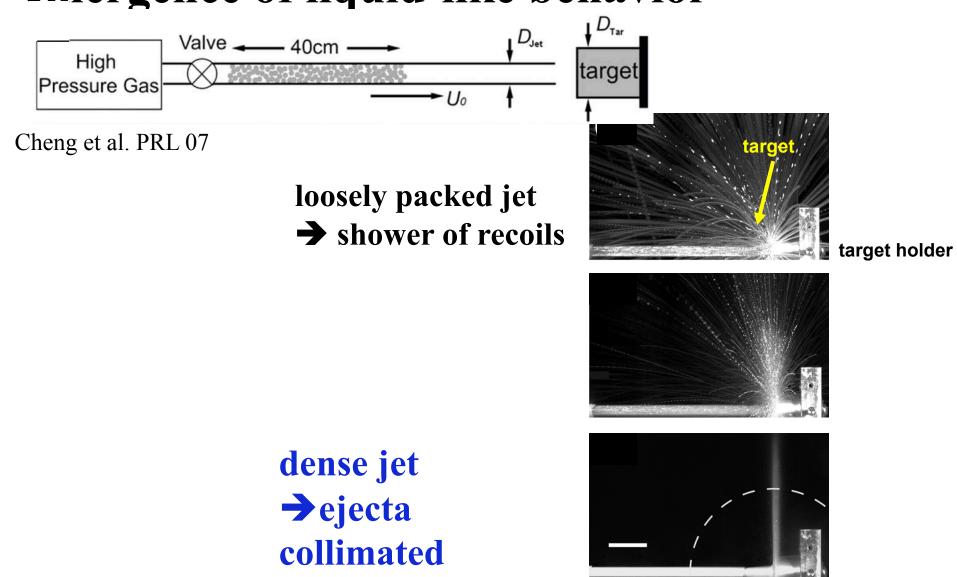
Impact can be surprising



4 mm diameter ethanol drop, impact speed = 4 m/s

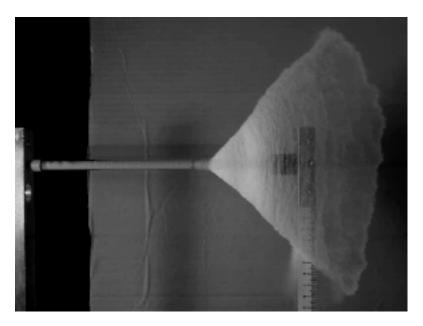
How does air create a splash?

Granular jet impact emergence of liquid-like behavior

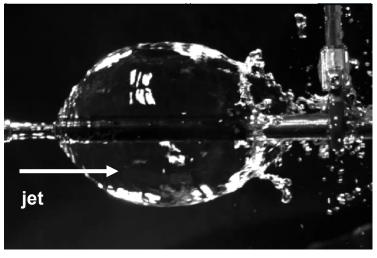


into thin sheet

Impact of thick jet onto small target



→ hollow conical sheet

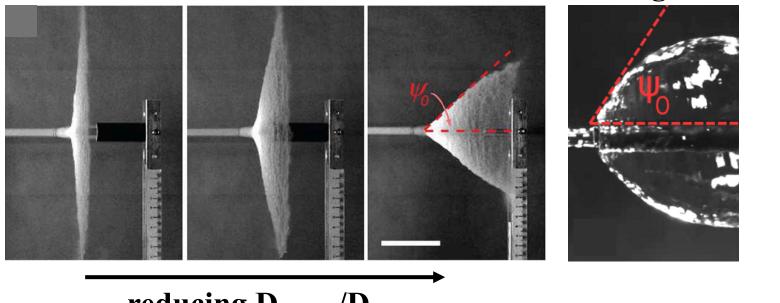


Water jet impact

→ hollow water bell

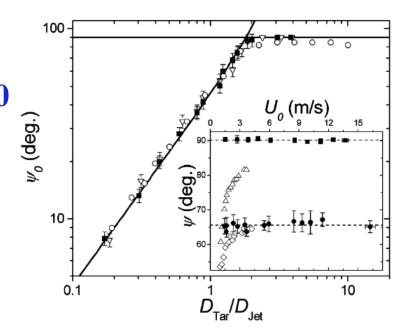
Granular "water bell"?

Ejecta sheet angle changes with D_{Target}/D_{Jet}



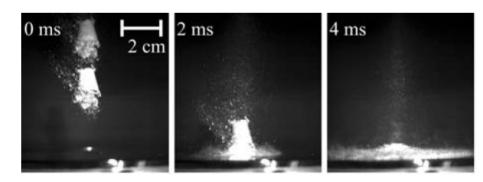
reducing D_{Target}/D_{Jet}

Granular ejecta angle ψ₀ agree numerically with values for water jet → liquid-like behavior



Context: emergence of collimated ejecta

 Formation of planetismals via collison of dust aggregates



ejecta collimated within 1°

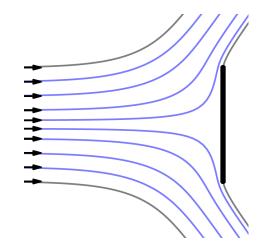


- Collimated ejecta from collision of gold ion jets at relativistic speeds
- -- Have been interpreted as evidence for a liquid quark-gluon phase

Teiser & Wurm, Mon. Not. R. Astron. Soc. 2009 Pozkanser, Voloshin, Ritter... 2008 APS Bonner prize talk Romatschke & Romatschke PRL 2007

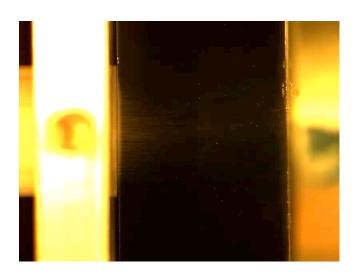
Interior motion different from water jet impact

impact of water jet continous motion no dead zone

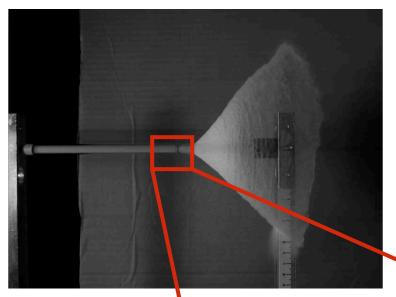


impact of granular jet flowing & static region

dead zone

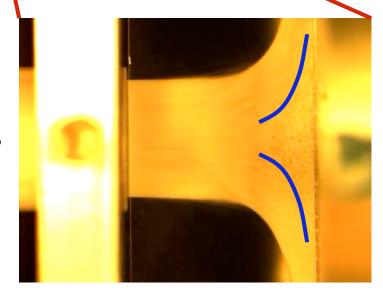


Liquid-like ejecta co-exists with solid-like interior



Ejecta agree quantitatively despite dissimilar interior

How?



Maybe high-density granular impact generically produces liquid-like ejecta, regardless of nature of interior state

Suppose we get rid of dead zone, would we still see collimated ejecta

Need numerics

the plan

- 1. Reproduce dead zone & ejecta sheet
- 2. Varying parameters to get rid of dead zone

Simulation

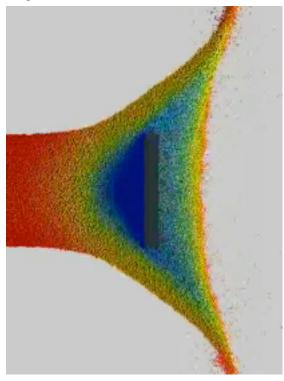
jet →

QuickTime™ and a decompressor are needed to see this picture. Minimal Physics Model
perfectly rigid grains
spheres
inelastic collisions
friction between grains
friction at target

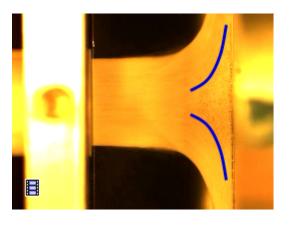
red = high speed blue = zero speed

Simulation vs. experiment

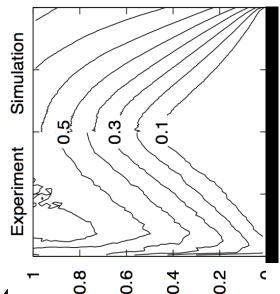
jet →



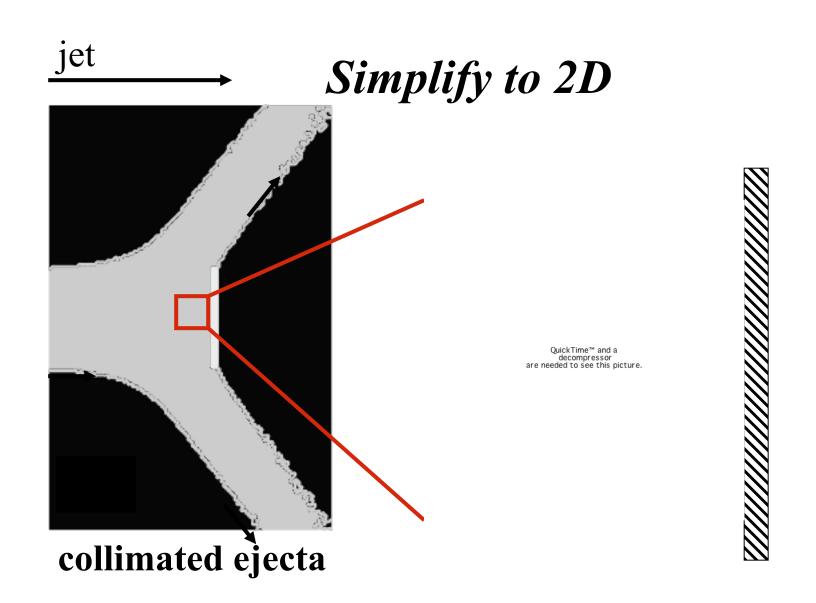
red = high speed
blue = zero speed



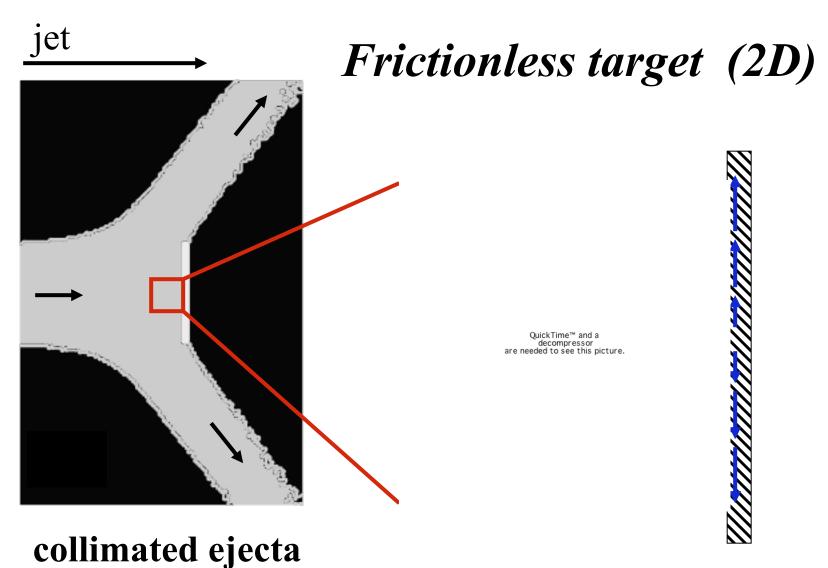
velocity contour comparison



Quantitative agreement height above target / target radius



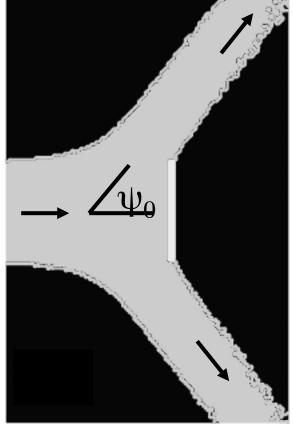
Same impact dynamics



liquid-like response without dead zone

jet





2D ejecta angle changes slightly

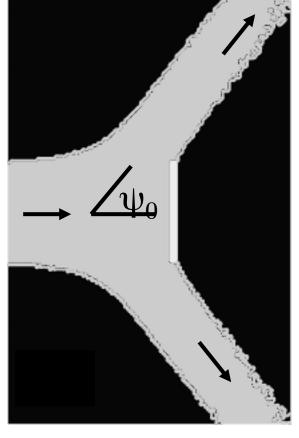
52° (with dead zone)

→ 47° (without deadzone)

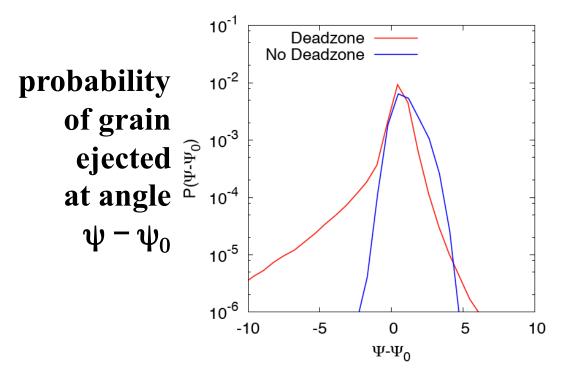
3D also slight change

jet

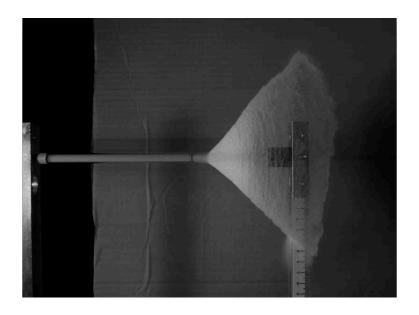
Collimated ejecta is generic



3D simulation



Why liquid-like ejecta?



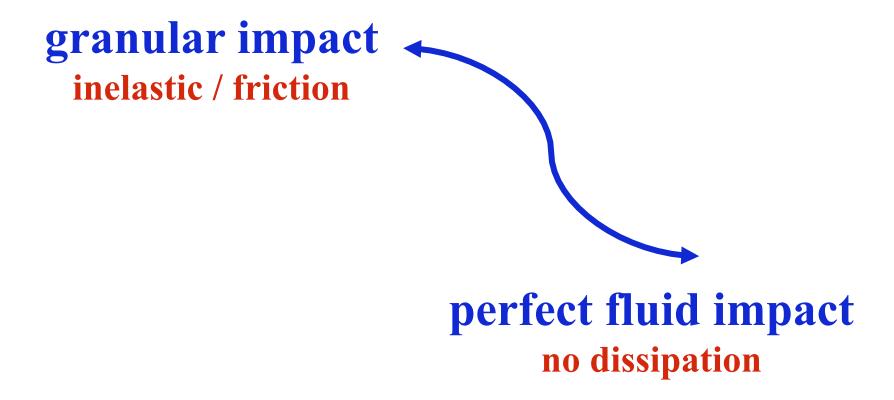
Granular & water jet impact controlled by same idealized limit of perfect fluid flow

Perfect fluid flow contains no information

Demonstrating connection between granular impact and perfect fluid flow

granular impact inelastic / friction no cohesion 2D simulation detailed perfect fluid impact quantitative no dissipation no surface tension comparison 2D exact solution

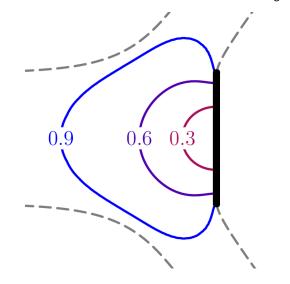
How good should the agreement be?

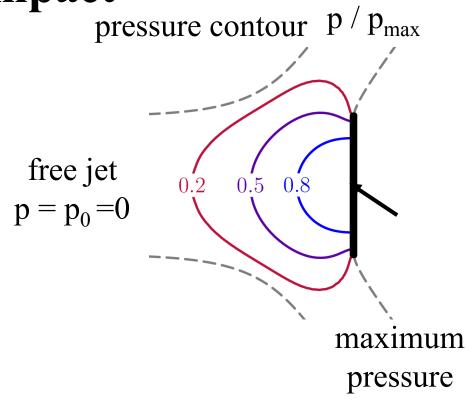


expect granular ejecta to move slower

Exact solution for perfect fluid impact

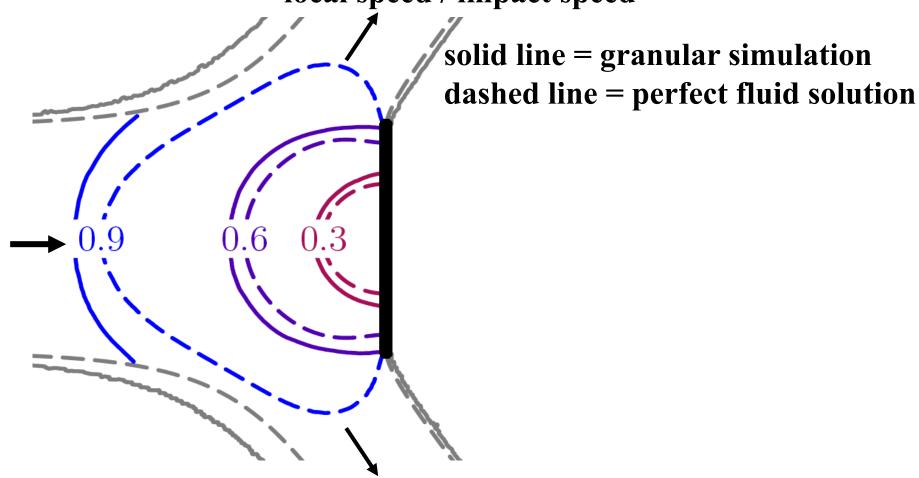
velocity contour u/U₀





Velocity contours

local speed / impact speed

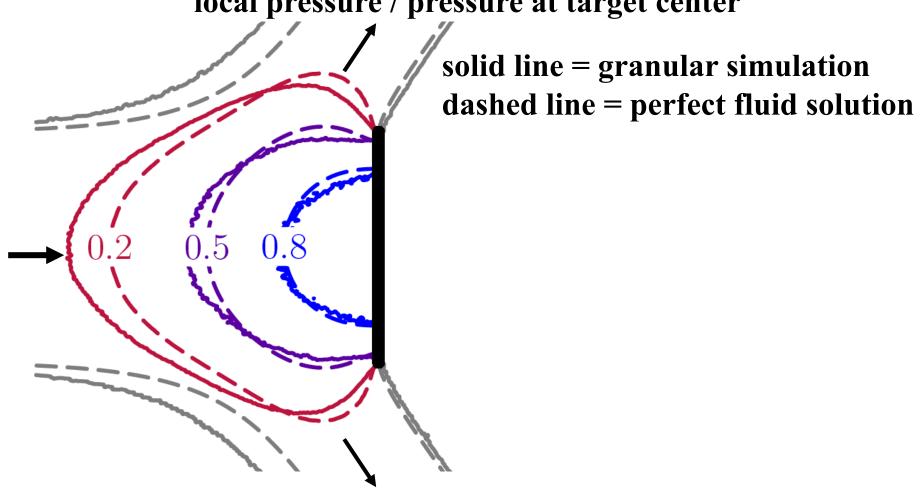


Inelasticity & friction

→ejecta wider than perfect fluid

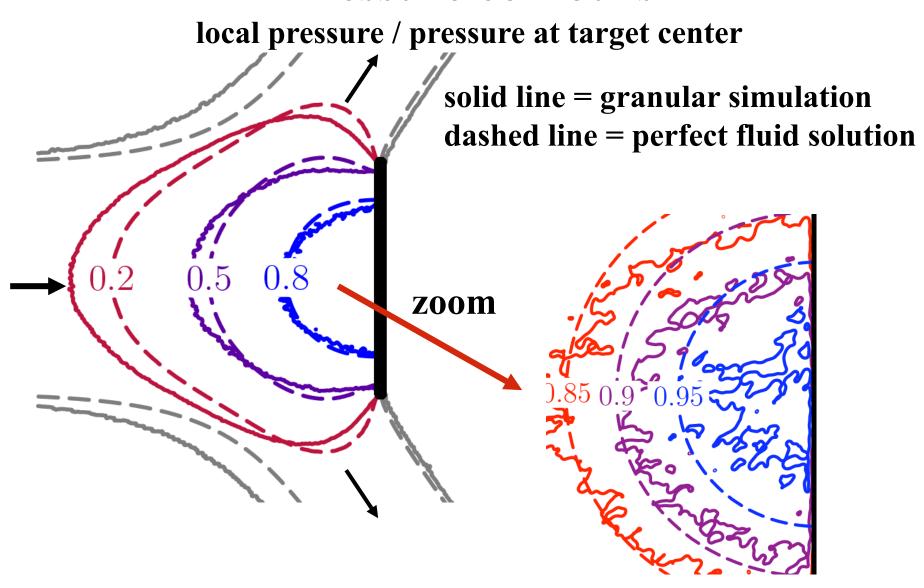
Pressure contours





Quantitative agreement

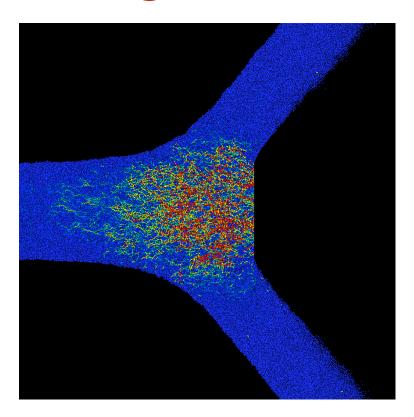
Pressure contours



Larger fluctuations

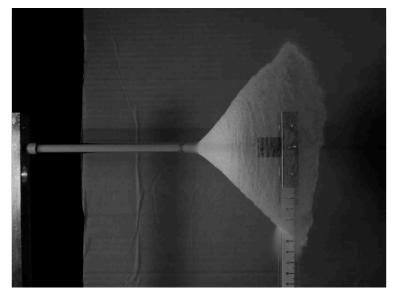
Distribution of compressive forces on grains

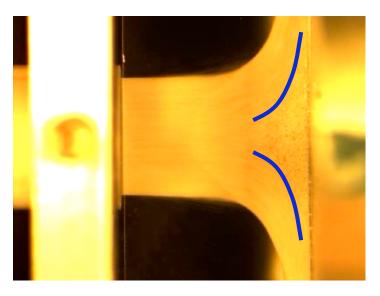
red = large blue = zero



force chains → fluctuations about perfect fluid behavior

Why liquid-like ejecta?





Generic outcome of high density impact

→ Collimated, liquid-like ejecta

With Nicholas Guttenberg, Jake Ellowitz,

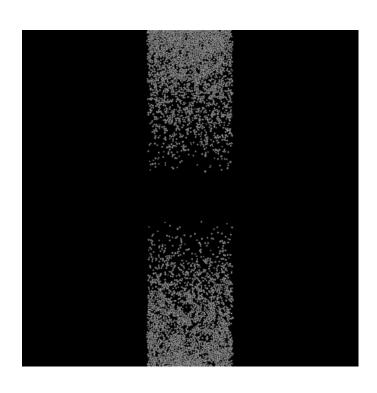
Herve Turlier, Sidney R. Nagel

Acknowledgements: Xiang Cheng, Efi Efrati, Heinrich M. Jaeger NSF-MRSEC, Keck Foundation, NSF-CBET

Thank you



What next?



Simulating dense granular impact is hard

kinetic regime (gas-like)

event-driven, hard-particle dynamics instantaneous collision exact momentum & energy conservation

★ vulnerable to numerical singularity (inelastic collapse)

standard toolkit

quasi-static, dense packing

discrete-element method elastic energy penalizing overlap

- **★** particles much softer than real materials
- ★ does not conserve energy exactly

Simulating dense granular impact is hard

kinetic regime (gas-like)

event-driven, hard-particle dynamics instantaneous collision exact momentum & energy conservation

★ vulnerable to numerical singularity (inelastic collapse)

dense flow conserving momentum & energy

quasi-static, dense packing

discrete-element method elastic energy penalizing overlap

- **★** particles much softer than real materials
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standard toolkit

How we simulate dense granular impact



kinetic regime (gas-like)
event-driven, hard-particle dynamics
instantaneous collision
exact momentum & energy conservation

★ vulnerable to numerical singularity (inelastic collapse)

Modified event-driven dynamics

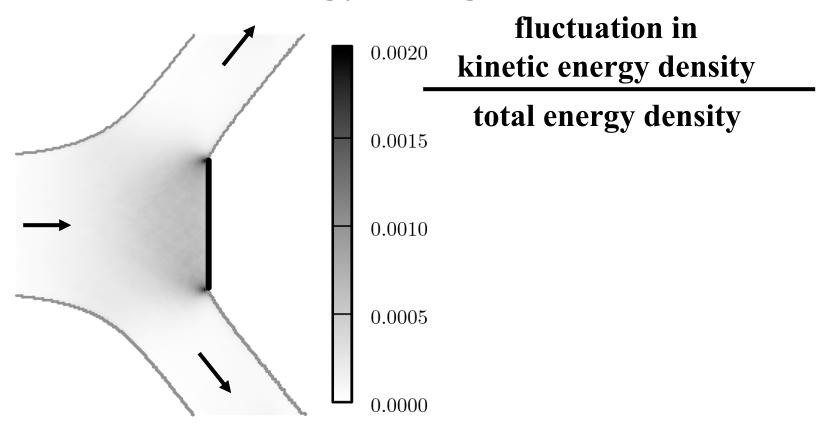
Evolve dynamics in fixed time interval Δt At time t find all particles that will overlap during Δt Avoid overlap by pretending the particles have collided at t Evolve collisions

Iterate until no overlap occurs in Δt Evolve to next time-step

McNamara, Flekk y & M al y PRE 2000

Guttenberg arXiv:1102.2483v1

Energy budget



Dominated by mean flow

Mass & momentum budget

Nearly incompressible flow velocity field

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{u}) = \mathbf{0} \implies \frac{\nabla \cdot \mathbf{u} = 0}{\text{constant & uniform density}}$$

High impact speed, large deceleration

- **→**neglect dissipation
- momentum transport has only inertia

pressure field
$$\rho \left(\frac{\partial}{\partial t} + \mathbf{u} \cdot \nabla \right) \mathbf{u} + \nabla p = 0$$

Perfect fluid flow

Only pressure gradient No shear stress

Pressure in granular impact

Define time interval τ << impact time-scale Define sample region (~ 5 particles wide) Sum impulses I_n experienced by each particle Define stress component as

$$\sigma_{ij} \propto \frac{1}{\tau} \sum_{n} (\mathbf{I}_n \cdot \hat{\mathbf{e}}_i)(\hat{\mathbf{r}}_n \cdot \hat{\mathbf{e}}_j), \quad i, j \in \{x, y\}$$

Define pressure as

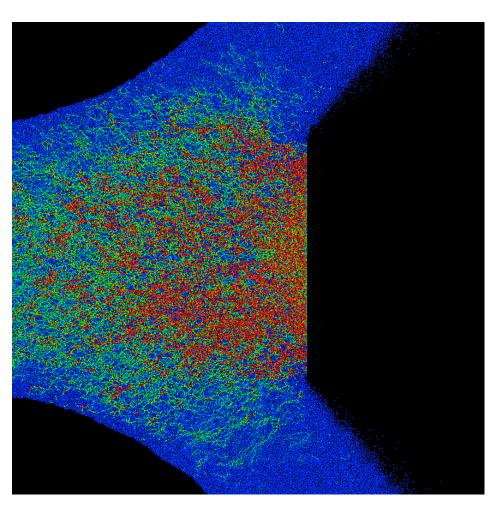
center-of-mass vector between colliding particles

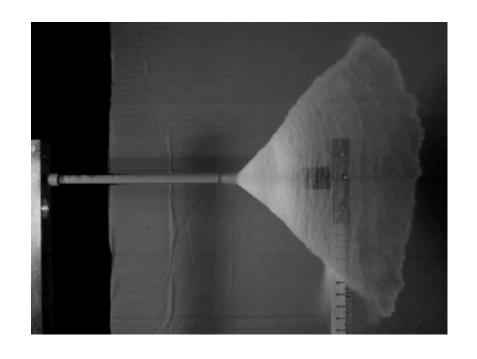
$$P = (|\sigma_{xx}| + |\sigma_{yy}|)/2$$

Average over T >> impact time-scale for contours

Distribution of compressive forces on grains with dead zone present

red = large blue = zero



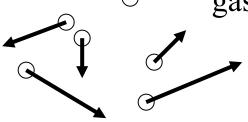


Granular jet impact

→ strongly coupled liquid?

When does liquid-like behavior emerge?

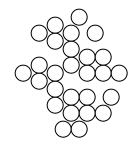
Emergence of liquid-like behavior gas child's view



• fills available volume

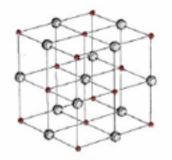
flows under shear





- fixed volume
- flows under shear

crystalline solid



- fixed volume
- resists shear

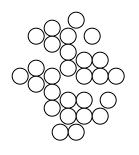
an intermediate state of matter

Emergence of liquid-like behavior gas traditional view

liquid

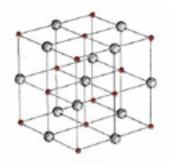


• flows under shear



- fixed volume
- flows under shear

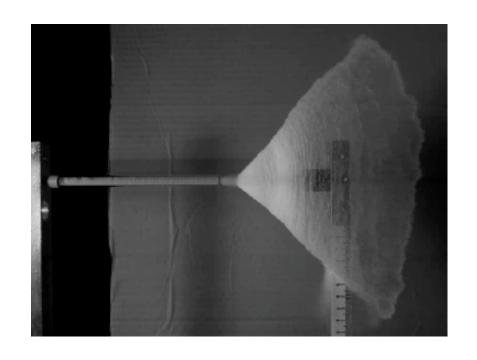
crystalline solid



- fixed volume
- resists shear

attractive interaction

liquid



Cohesion between particles effectively zero

Emergence of liquid-like behavior modern view

In a liquid spatial arrangement of molecules is controlled

- mostly by strong repulsion between nearby neighbors
 - → can model molecules as hard spheres
- attraction perturbs microscopic spatial structure
 - → can model attraction by confinement
- → computation scheme yielding structure & equation of state

confine hard spheres

liquid

Velocity along centerline

Speed along center of jet

